# Strange Nucleon Form Factors from *ep* and *vp* Elastic Scattering

A combined analysis of HAPPEx,  $G^0$ , and BNL E734 data

Stephen Pate New Mexico State University

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## Outline

- Parity-violating electron-nucleon (PVeN) elastic scattering
- Neutrino-proton elastic scattering
- How the use of neutrino-proton elastic scattering data with PVeN data permits the extraction of the strange vector and axial form factors
- The results of a combined analysis of BNL E734 vp data with the HAPPEX and  $G^0$  forward PVeN data begin to show the  $Q^2$ -dependence of the strange axial form factor for  $Q^2 = 0.45 1.0 \text{ GeV}^2$ .
- These results greatly restrict the possible configurations of strange quarks in the nucleon.
- How these results will aid the determination of the  $Q^2$ -dependence of the strange vector form factors

# Features of parity-violating forward-scattering *ep* data

- measures linear combination of form factors of interest
- axial terms are doubly suppressed

 $\rightarrow (1 - 4\sin^2\theta_W) \sim 0.075$ 

- $\rightarrow$  kinematic factor  $\varepsilon' \sim 0$  at forward angles
- significant radiative corrections exist, especially in the axial term

parity-violating data at forward angles are mostly sensitive to the strange electric and magnetic form factors

## Full Expression for the PV ep Asymmetry

For a hydrogen target, the asymmetry as a linear combination of  $G_E^s$ ,  $G_M^s$ ,  $G_A^{CC}$  and  $G_A^s$  is:

$$A^{p} = A_{0}^{p} + A_{E}^{p}G_{E}^{s} + A_{M}^{p}G_{M}^{s} + A_{AIV}^{p}G_{A}^{cC} + A_{A}^{p}G_{A}^{s}$$
where  $A_{0}^{p} = -K^{p} \begin{cases} \left(1 - 4\sin^{2}\theta_{W}\right)\left(1 + R_{V}^{p}\right)\left(\varepsilon G_{E}^{p^{2}} + \tau G_{M}^{p^{2}}\right)\right) \\ -\left(1 + R_{V}^{n}\right)\left(\varepsilon G_{E}^{p}G_{E}^{n} + \tau G_{M}^{p}G_{M}^{n}\right) \\ -\varepsilon'G_{M}^{p}\left(1 - 4\sin^{2}\theta_{W}\right)\left[\sqrt{3}R_{A}^{T=0}G_{A}^{8}\right] \end{cases}$ 

$$A_{E}^{p} = K^{p} \left\{\varepsilon G_{E}^{p}\left(1 + R_{V}^{0}\right)\right\}$$

$$A_{M}^{p} = K^{p} \left\{\tau G_{M}^{p}\left(1 + R_{V}^{0}\right)\right\}$$

$$A_{A}^{p} = K^{p} \left\{\varepsilon' G_{M}^{p}\left(1 - 4\sin^{2}\theta_{W}\right)\left(1 + R_{A}^{T=1}\right)\right\}$$

$$A_{A}^{p} = K^{p} \left\{\varepsilon' G_{M}^{p}\left(1 - 4\sin^{2}\theta_{W}\right)\left(1 + R_{A}^{0}\right)\right\}$$

$$K^{p} = \frac{G_{F}Q^{2}}{4\pi\sqrt{2}\alpha} \frac{1}{\varepsilon G_{E}^{p^{2}} + \tau G_{M}^{p^{2}}}$$
Note suppression of axial terms by  $(1 - 4\sin^{2}\theta_{W})$  and

ε'.

#### Things known and unknown in the PV ep Asymmetry

$$G_{E,M}^{p,n}$$
 = Kelly parametrization [PRC 70 (2004) 068202]

with  $G^0$  uncertainties [http://www.npl.uiuc.edu/exp/G0/Forward]

$$G_{A}^{CC} = \frac{g_{A}}{\left(1 + Q^{2}/M_{A}^{2}\right)^{2}} \qquad G_{A}^{8} = \frac{1}{2\sqrt{3}} \frac{\left(3F - D\right)}{\left(1 + Q^{2}/M_{A}^{2}\right)^{2}} M_{A} = 1.001 \pm 0.020 \text{ GeV} \left[\text{Budd, Bodek and Arrington : hep - ex/0308005 and 0410055}\right] g_{A} = 1.2695 \pm 0.0029 \left[\text{Particle Data Group 2005}\right] 3F - D = 0.585 \pm 0.025 \left[\text{Goto et al. PRD 62 (2000) 034017}\right] \left[\text{use of } 3F - D \text{ implies use of flavor - SU(3), but } G_{A}^{8} \text{ is suppressed by } \varepsilon' \text{ and } \left(1 - 4\sin^{2}\theta_{W}\right)\right]$$

The *R*'s are radiative corrections calculated at  $Q^2 = 0$  in the formalism of Zhu et al. [PRD 62 (2000) 033008]. The  $Q^2$  - dependence is unknown, and so we have assigned a 100% uncertainty to the values.

 $R_{\rm V}^{p} = -0.045 \qquad R_{\rm V}^{n} = -0.012 \qquad R_{\rm V}^{0} = -0.012$  $R_{\rm A}^{T=1} = -0.173 \qquad R_{\rm A}^{T=0} = -0.253 \qquad R_{\rm A}^{0} = -0.552$ [from evaluation of Arvieux*et al.*, to be published]

### Features of elastic vp data

- measures quadratic combination of form factors of interest
- axial terms are dominant at low  $Q^2$

$$\frac{d\sigma}{dQ^2} \left( vp \rightarrow vp \right) \xrightarrow{Q^2 \rightarrow 0} \frac{G_F^2}{128\pi} \frac{M_p^2}{E_v^2} \left[ \left( -G_A^u + G_A^d + G_A^s \right)^2 + \left( 1 - 4\sin^2 \theta_W \right)^2 \right]$$

• radiative corrections are insignificant

[Marciano and Sirlin, PRD 22 (1980) 2695]

meutrino data are mostly sensitive to the strange axial form factor

#### Elastic NC neutrino-proton cross sections

$$\frac{d\sigma}{dQ^2} \left( vp \rightarrow vp \right) = \frac{G_F^2}{2\pi} \frac{Q^2}{E_v^2} \left( A \pm BW + CW^2 \right) + v$$

$$W = 4\left(E_{v}/M_{p} - \tau\right) \qquad \tau = Q^{2}/4M_{p}^{2}$$
$$A = \frac{1}{4}\left[\left(G_{A}^{Z}\right)^{2}\left(1 + \tau\right) - \left(\left(F_{1}^{Z}\right)^{2} - \tau\left(F_{2}^{Z}\right)^{2}\right)\left(1 - \tau\right) + 4\tau F_{1}^{Z}F_{2}^{Z}\right]$$

$$B = -\frac{1}{4}G_{A}^{Z}\left(F_{1}^{Z} + F_{2}^{Z}\right)$$
$$C = \frac{1}{64\tau}\left[\left(G_{A}^{Z}\right)^{2} + \left(F_{1}^{Z}\right)^{2} + \tau\left(F_{2}^{Z}\right)^{2}\right]$$

Dependence on strange form factors is buried in the weak (Z) form factors.

## The BNL E734 Experiment

- performed in mid-1980's
- measured neutrino- and antineutrino-proton elastic scattering
- used wide band neutrino and anti-neutrino beams of  $\langle E_{v} \rangle = 1.25 \text{ GeV}$
- covered the range  $0.45 < Q^2 < 1.05 \text{ GeV}^2$
- large liquid-scintillator target-detector system
- still the **only** elastic neutrino-proton cross section data available

#### E734 Results Uncertainties shown are total (stat and sys). Correlation coefficient arises from systematic errors.

$Q^2$	$\frac{d\sigma/dQ^2(\nu p)}{(\Gamma - V)^2}$	$\frac{d\sigma/dQ^2(\bar{\nu}p)}{(\bar{\nu}p)^2}$	correlation
$(\text{GeV})^2$	$(\text{Im/GeV})^2$	$(\text{Im/GeV})^2$	coefficient
0.45	$0.165 \pm 0.033$	$0.0756 \pm 0.0164$	0.134
0.55	$0.109 \pm 0.017$	$0.0426 \pm 0.0062$	0.256
0.65	$0.0803 \pm 0.0120$	$0.0283 \pm 0.0037$	0.294
0.75	$0.0657 \pm 0.0098$	$0.0184 \pm 0.0027$	0.261
0.85	$0.0447 \pm 0.0092$	$0.0129 \pm 0.0022$	0.163
0.95	$0.0294 \pm 0.0074$	$0.0108 \pm 0.0022$	0.116
1.05	$0.0205 \pm 0.0062$	$0.0101 \pm 0.0027$	0.071

#### Combination of the ep and vp data sets

We use PV *ep* data in the same range of  $Q^2$  as the E734 experiment.

• The original HAPPEx measurement:  $Q^2 = 0.477 \text{ GeV}^2$ [PLB 509 (2001) 211 and PRC 69 (2004) 065501]

• The recent  $G^0$  data covering the range  $0.1 < Q^2 < 1.0 \text{ GeV}^2$ [PRL 95 (2005) 092001]



### Combination of the ep and vp data sets

Since the neutrino data are quadratic in the form factors, then there will be in general <u>two solutions</u> when these data sets are combined.



Fortunately, the two solutions are very distinct from each other, and other available data can select the correct physical solution.

#### **General Features of the two Solutions**

	Solution 1	Solution 2
$G_E{}^s$	Consistent with zero (with large uncertainty)	Large and positive
$G_M^{s}$	Consistent with zero (with large uncertainty)	Large and negative
$G_{A}{}^{s}$	Small and negative	Large and positive

There are three strong reasons to prefer **Solution 1**:

- $G_A^s$  in Solution 2 is inconsistent with DIS estimates for  $\Delta s$
- $G_M{}^s$  in Solution 2 is inconsistent with the combined SAMPLE/PVA4/HAPPEx result of  $G_M{}^s = \sim +0.6$  at  $Q^2 = 0.1 \text{ GeV}^2$
- $G_E{}^s$  in Solution 2 is inconsistent with the idea that  $G_E{}^s$  should be small, and conflicts with expectation from recent  $G^0$  data that  $G_E{}^s$  may be negative near  $Q^2 = 0.3 \text{ GeV}^2$

I only present Solution 1 in what follows.











• G0 & E734 [to be published]

 $Q^2$ -dependence suggests  $\Delta s < 0$  !

○ HAPPEx & E734 [Pate, PRL 92 (2004) 082002]



- G0 & E734 [to be published]
- HAPPEx & E734 [Pate, PRL 92 (2004) 082002]

Recent calculation by Silva, Kim, Urbano, and Goeke (hep-ph/0509281 and Phys. Rev. D 72 (2005) 094011) based on chiral quark-soliton model is in rough agreement with the data.



An, Riska and Zou, hep-ph/0511223; Riska and Zou, nucl-th/0512102.

#### Strange Vector Form Factors: Using *ep* and *vp* data

The international program of PV ep measurements will completely resolve the strange vector form factors at only three  $Q^2$  points: 0.1, 0.23 and 0.63 GeV<sup>2</sup>.

At many other points in the range  $0.038 < Q^2 < 1.0 \text{ GeV}^2$ , we have PV physics asymmetries that represent linear combinations of the vector, axial and anapole form factors. These can be used to constrain fits that seek to understand the  $Q^2$ -dependence of these form factors.

As has just been demonstrated, a combination of the forwardscattering PV *ep* data with elastic *vp* data provides several more points where the vector form factors are resolved, and with reasonable error bars: 0.63, 0.79, and 0.99 GeV<sup>2</sup>. These can already provide powerful constraints on global fits.



#### But better vp data are needed!

The E734 data have insufficient precision and too narrow a  $Q^2$  range to achieve the full potential of this physics program. Better neutrino data is needed, with smaller uncertainties and points nearer  $Q^2 = 0$ , to fulfill the potential of this analysis method.

A detailed understanding of the  $Q^2$ -dependence of these form factors will not be possible until a more dense set of resolved data points are available. FINeSSE can provide these additional data points.

