Overview of Neutrino Experiments and Facilities

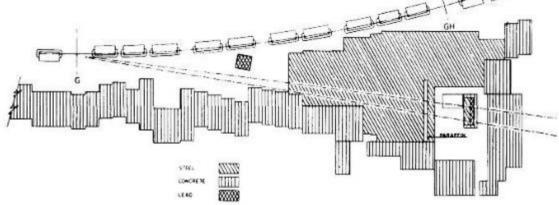
Deborah Harris Workshop on Intersections of Nuclear Physics with Neutrinos and Electrons Newport News, Virginia May 4-5, 2006

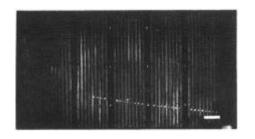
Outline

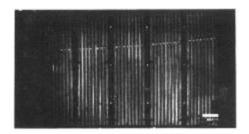
- Brief History of Neutrino Beams and Experiments
- What you can and cannot measure with neutrino scattering events
- What beamlines and detectors are currently (or soon to be) in operation

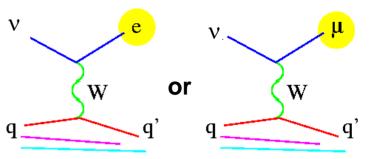
1st Neutrino Beam Experiment

• 1960: Used to determine neutrinos have lepton flavor like muon and electron Nobel citation: "for the neutrino beam method and the demonstration of the doublet structure of the leptons through the discovery of the muon-neutrino"







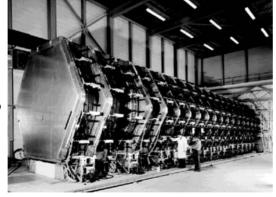


- If $\mu^+ \rightarrow e^+ \nu \overline{\nu}$, why not $\mu^+ \rightarrow e^+ \gamma$?
- 5BeV protons on Be Target (3.5x10¹⁷ of them)
- $\pi^+ \rightarrow \mu^+ \nu_{\mu}$
- 21m decay region
- 13.5m Fe Shield, 1 Ton Detector
- 34 single-µ events
- 5 background
- NO e-like events!
- PRL: 1960, NP: 1988
- Lederman, Swartz, Steinberger Brookhaven Nat'l Lab

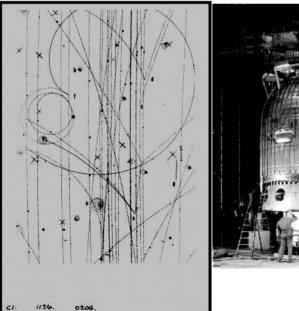
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Neutrino Experiments

- Beamline Progression: Add focusing horns
- Detector Progression: look at more than just "is there a muon or not"
- Sixties and Seventies: Measurements of nucleon structure
 - CDHS, CHARM, CHARMII, BEBC



Trying to see what else gets produced with neutrinos...Resonances Kaons, Di-muons (evidence of charm)

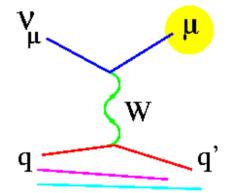


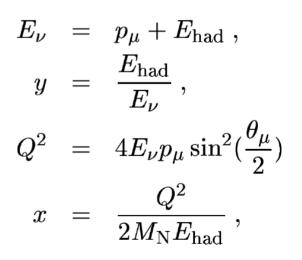
s,d

С

What is and isn't measured in v Scattering: Charged Currents

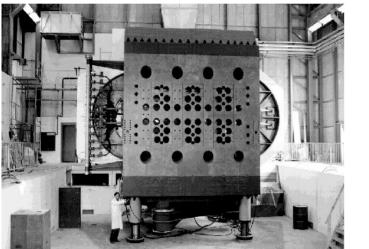
- Outgoing Lepton
 - Momentum (p_{μ})
 - Charge (maybe: determines if v or vbar)
 - Flavor
 - Angle
- Other outgoing particles:
 - Total visible energy
 - If Particle Identification: could add rest masses too
- What is known *a priori*: v direction



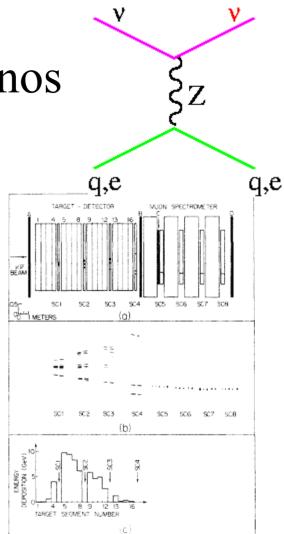


Neutral Currents with Neutrinos

1973: Neutral Currents first seen! - Gargamelle at CERN, E1A at Fermilab







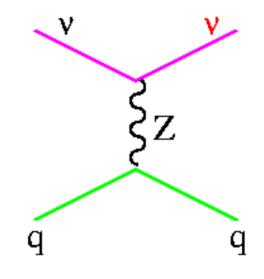
• 1980: Higher energy beams: higher Q², more events, better precision on $\sin^2\theta_{m}$ $R^{\nu(\overline{\nu})} = \frac{\sigma_{NC}^{\nu(\overline{\nu})}}{\sigma_{CC}^{\nu(\overline{\nu})}} = \rho^2 \left(\frac{1}{2} - \sin^2\theta_W + \frac{5}{9}\sin^4\theta_W \left(1 + \frac{\sigma_{CC}^{\nu(\nu)}}{\sigma_{CC}^{\nu(\overline{\nu})}}\right)\right)$

- CHARM, CCFR
$$R^{\nu(\bar{l})}$$

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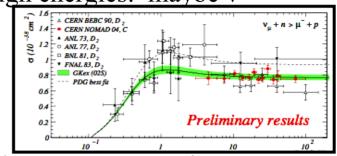
What is and isn't measured in v Scattering: Neutral Currents

- Outgoing Lepton
 - Momentum (p_{μ})
 - Charge (maybe: determines if v or vbar)
 - Flavor – Angle
- Other outgoing particles:
 - Total visible energy
 - If PID: could add rest masses too
- What is known *a priori*: v direction
- And if you have really pure beams of only one charge of pion, can know if it's a neutrino vs antineutrino...

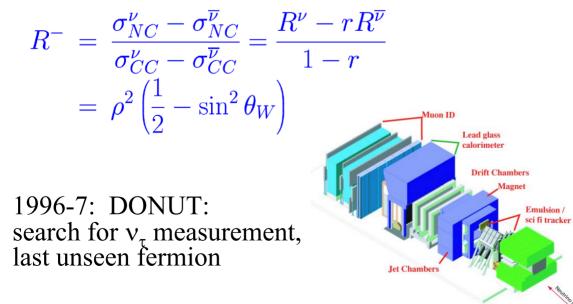


Neutrinos in the Nineties

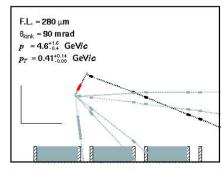
- Searches for v oscillations at short distances, high energies: maybe v matrix is like quark matrix...
 - CHORUS, NOMAD
 - Structure functions and cross sections
 - R. Petti, NOMAD Nuint05, ¹²C target



• 1996-7: NuTeV: High intenstiy sign-selected beam: pure neutrino or antineutrino beam for $\sin^2\theta_W$



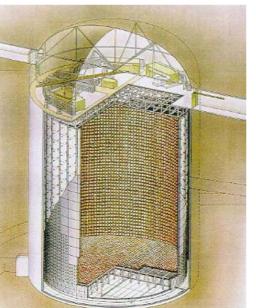


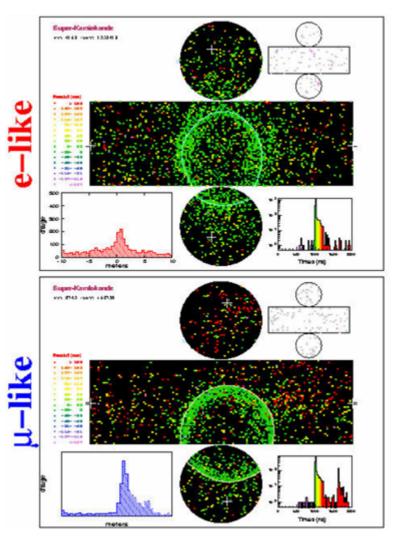


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New Direction for v Experiments

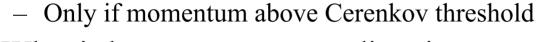
- 1998: "Smoking Gun Evidence" for Neutrino Oscillations in Atmospheric neutrinos
- 300km/GeV becomes an interesting place to look
- Water Cerenkov detector takes center stage
- 1993: LSND signal seen, implying 1km/GeV is another signal region



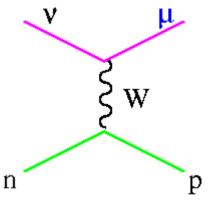


What is and isn't measured in v Scattering: Water Cerenkov Detectors

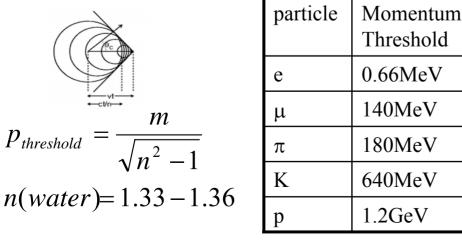
- Outgoing Lepton
 - Momentum (p_{μ})
 - Flavor
 - Angle
- Other outgoing particles



• What is known *a priori*: v direction



If you assume Quasi-Elastic Event:



$$E_{\nu} = \frac{m_N E_{\mu} - m_{\mu}^2/2}{m_N - E_{\mu} + p_{\mu} \cos \theta_{\mu}}$$
$$Q^2 = 4E_{\nu} p_{\mu} \sin^2(\frac{\theta_{\mu}}{2})$$

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Sea Change in Neutrino Experiments

Medium Energy NuMI Beam Tune

0 mrad

11

rates for L = 810 km

30 25

20

15 10

- Now: Lower and Lower Energies
 - K2K, MINOS: v_{μ} disappearance over 250,735km
 - OPERA: $v_{\mu} \rightarrow v_{\tau}$ over 732km

x 10⁹

0.35

0.3

0.25

0.2

0.15

0.1

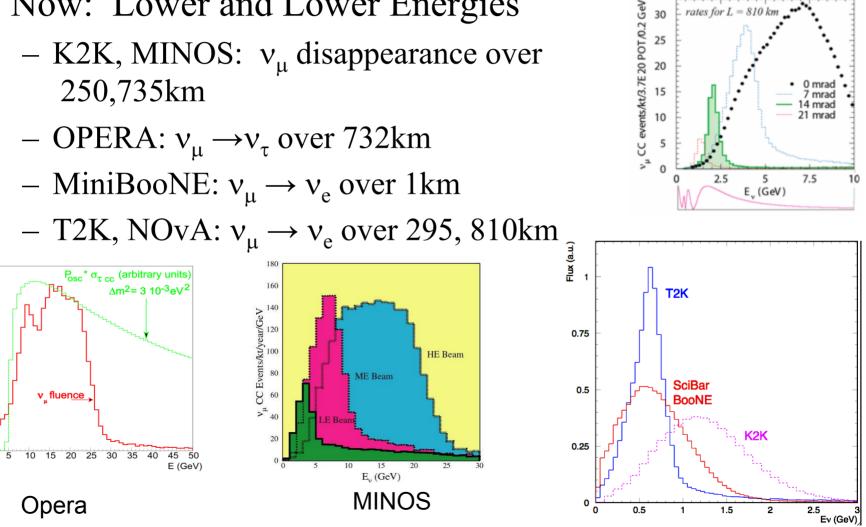
0.05

0 0

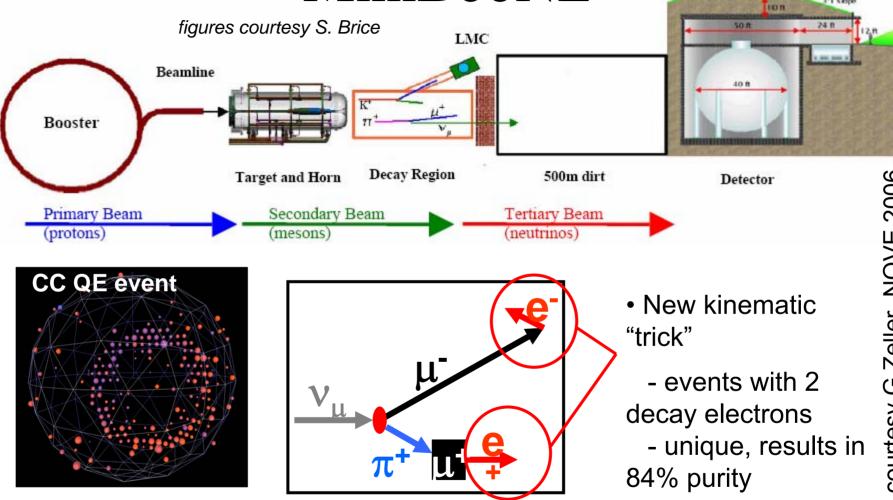
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 $m^2 j^l$ 0.4

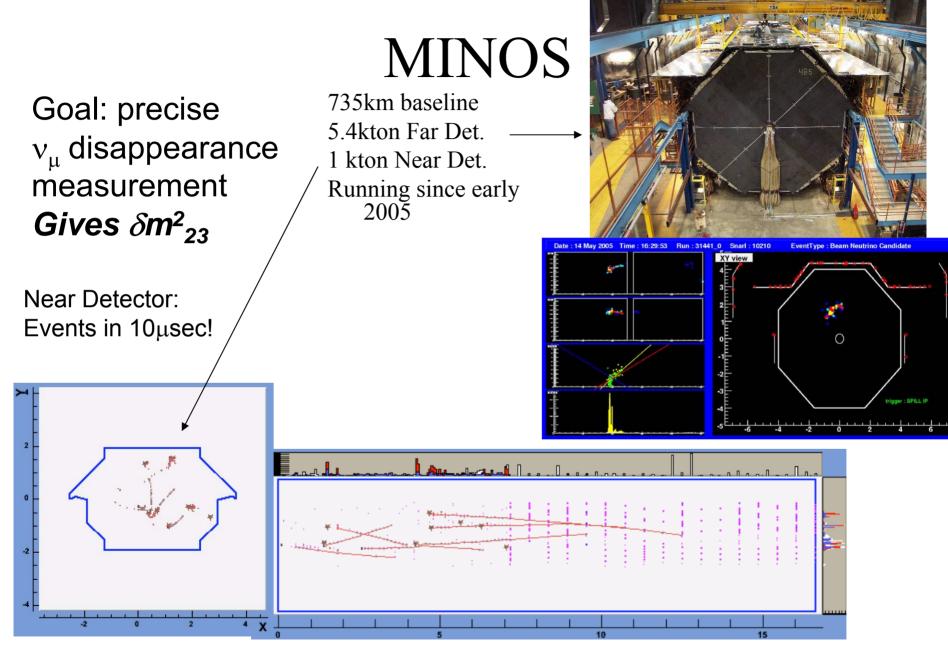
CoV.



MiniBooNE



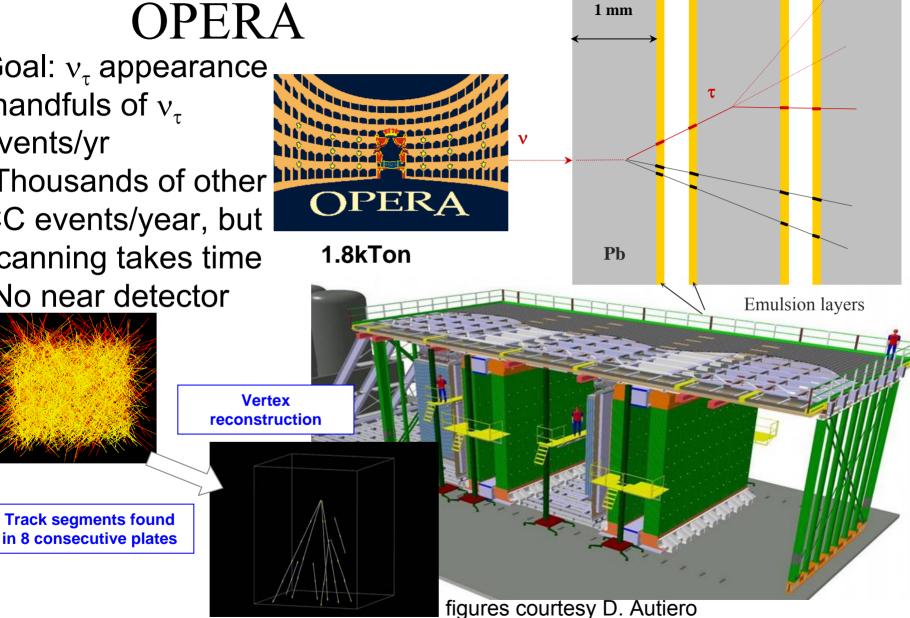
1-1 slope



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Goal: v_{τ} appearance •handfuls of v_{τ} events/yr

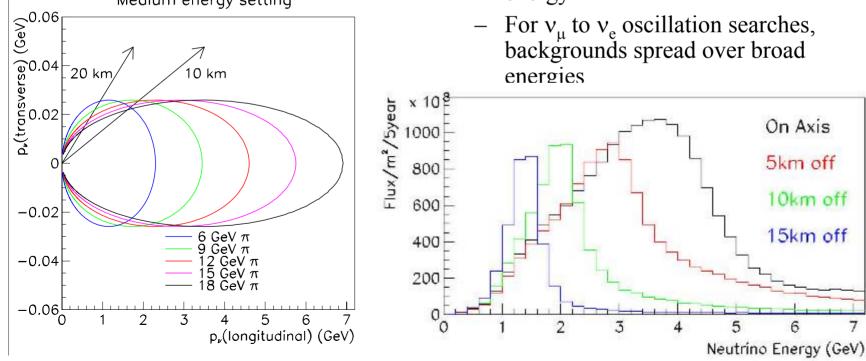
•Thousands of other CC events/year, but scanning takes time •No near detector



Off Axis Strategy



- Fewer total number of neutrino events
- More at one narrow region of energy
- For v_{μ} to v_{e} oscillation searches, backgrounds spread over broad energies

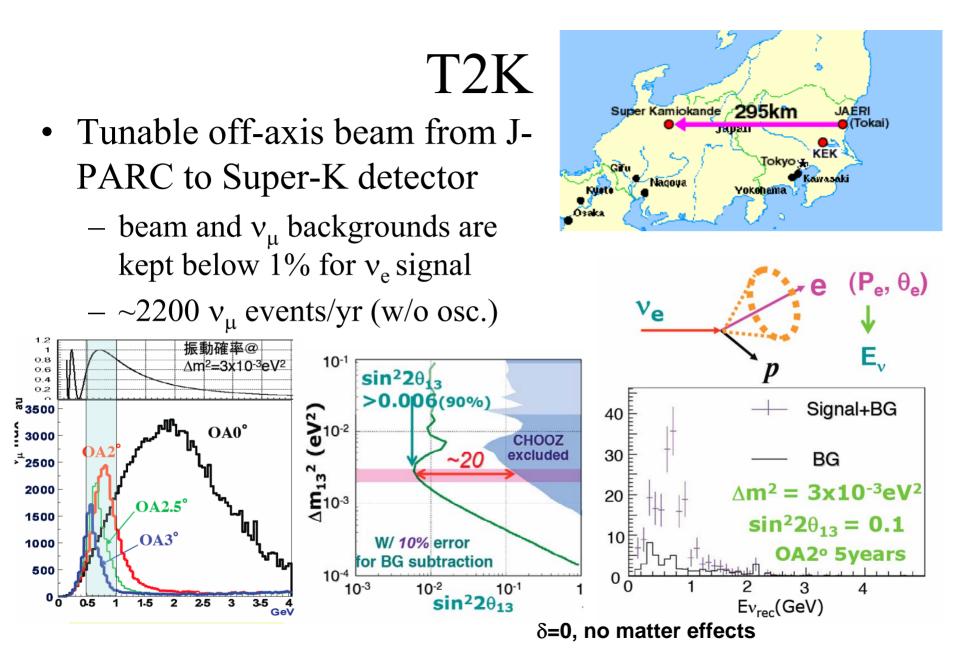


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______20 km

Medium energy setting

, 10 km



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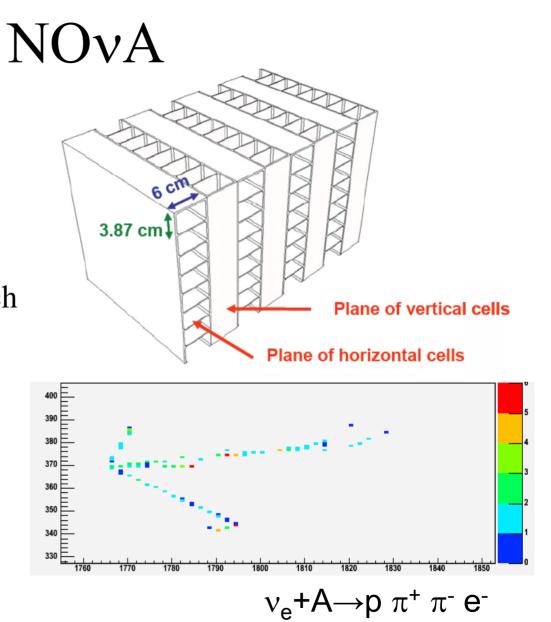
Use Existing NuMI beamline

- Build new 30kTon Scintillator Detector
- 810km baseline-compromise between reach in θ_{13} and matter effects
- Near detector: similar granularity, off-axis

Goal:

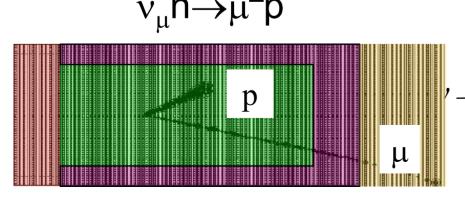
 ν_{e} appearance In ν_{μ} beam

figures courtesy J. Cooper



MINERvA

- **Dedicated Neutrino Interaction Experiment!**
- Fine-grained scintillator detector
- Surrounded by Calorimetry
- Nuclear targets upstream
- Focus on exclusive final state reconstruction and energy containment
- **Right in front of MINOS** near detector (muon spectrometer)
- To run in NuMI Beamline (Fermilab)



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 $v_{\mu}p \rightarrow v_{\mu}\pi^{0}p$

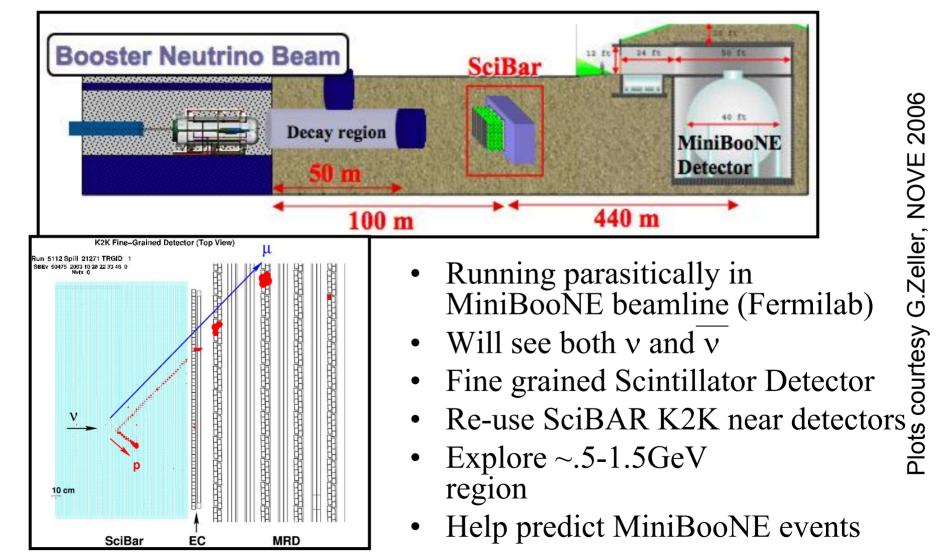
igures courtesy K.McFarland nuclear targets active detector

ECAL

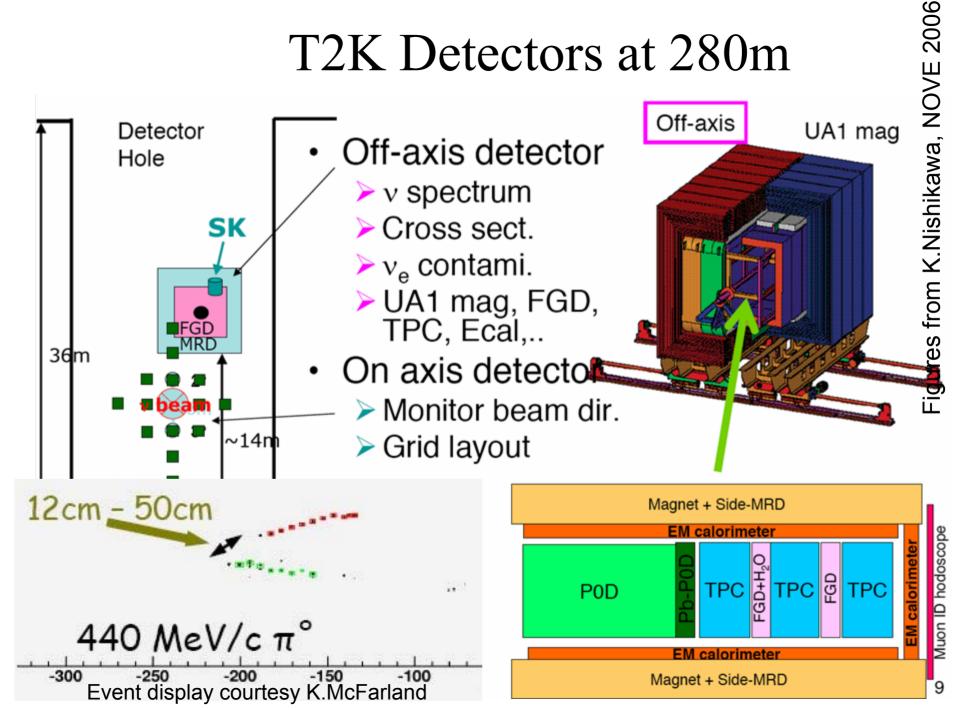
HCAL

18

SciBooNE

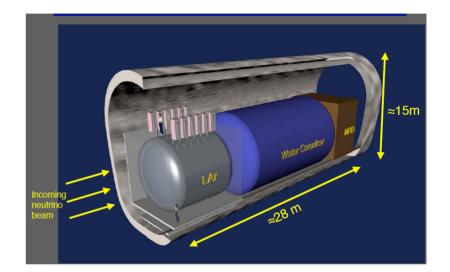


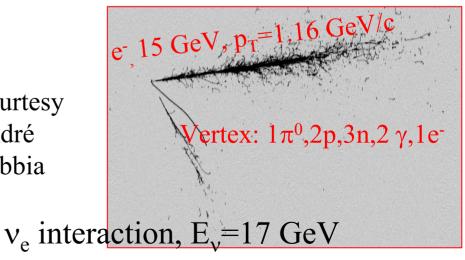
T2K Detectors at 280m

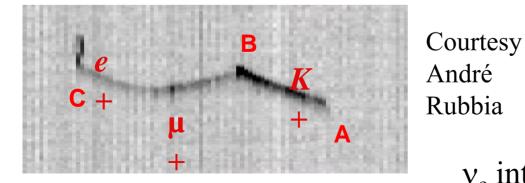


T2K Detectors at 2km

- Liquid Argon and Water Cerenkov
- Liquid Argon: electronic bubble chamber
 - Particle ID by dE/dx
 - Particle mometum by range
 - No magnetic field...

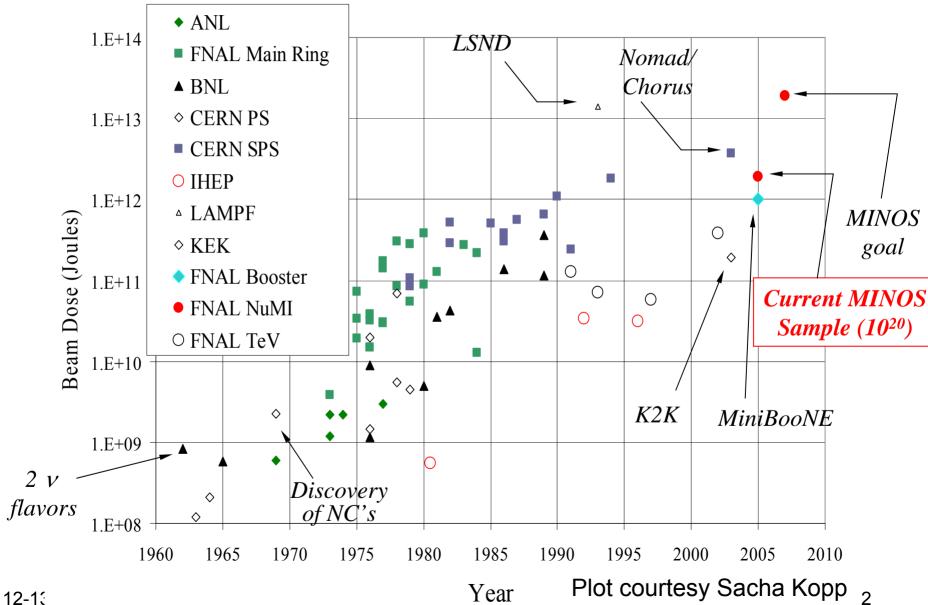






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Integrated proton power vs time...



Conclusions

- Neutrino Beamlines are getting more and more intense, but lower and lower energies.
- Neutrino detector physics is returning to its roots
 - Lower energy neutrinos needed for oscillation studies
 - Finer and finer grained detectors needed to see
 - $\nu_{\mu} \rightarrow \nu_{\tau}$
 - $v_{\mu} \rightarrow v_{e}$
- Many "tricks" up our sleeves
 - Changing beamline configurations in one run
 - Putting in several nuclear targets simultaneously
 - Measuring neutrino cross sections in these energy ranges precisely!
 - Taking advantage of really intense beamlines
 - use relatively light (very fine-grained) detectors