## High Strangeness Dibaryon Search with STAR Data

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## Outline

VIntroduction
छ $\mathrm{V} \Omega$ dibaryon
■ Two-particle correlation function
$-P \Omega$ correlation function
(ISummary and Outlook

## Introduction (1)

(V) Standard Model: Baryons - 3 quarks and Mesons - pair of quark-antiquark

I] 1977: within Quark Bag Model, Jaffe predicted H-dibaryon made of six quarks (uuddss) (Phys. Rev. Lett. 38,195 (1977); 38, 617(E)(1977))

I] Exotic hadrons - long standing challenge in hadron physics

Tetraquark
Meson-Meson molecule

Hexaquark
Baryon-Baryon molecule


## V Observation of exotic states @ WASA-at-COSY, Belle, LHCb




$\square$ Multi-quark states or molecular states?

Phys. Rev. Lett. 115 (2015) 072001
Phys. Rev. Lett. 112 (2014) 222002
Phys. Rev. Lett. 106 (2011) 242302

## Exotics in Strangeness Sector

Quark content, decay modes and mass of exotic states in strangeness sector:

| particle | Mass (MeV) | Quark composition | Decay mode |
| :---: | :---: | :---: | :---: |
| $\mathrm{f}_{0}$ | 980 | $q \bar{q} \mathrm{~s} \bar{s}$ | $\pi \pi$ |
| $\mathrm{a}_{0}$ | 980 | $q \bar{q} s \bar{s}$ | $\pi \eta$ |
| K(1460) | 1460 | $q \bar{q} q \bar{s}$ | K $\pi \pi$ |
| $\Lambda(1405)$ | 1405 | qqq s $\bar{q}$ | $\pi \Sigma$ |
| $\Theta^{+}(1530)$ | 1530 | q9q ${ }^{\text {¢ }}$ s | KN |
| H | 2245 | uuddss | \ |
| $N \Omega$ | 2573 | qqqsss |  |
| E $\underbrace{\square}$ | 2627 | qqissss | $\Lambda \Xi$ |
| $\Omega \Omega$ | 3228 | SSSSSS | $\Lambda \mathbf{K}^{-}+\Lambda \mathbf{K}^{-}$ |

Phys. Rev. C 84 (2011) 064910, Phys. Rev. C 83 (2011) 015202
■ Recent results on H -dibaryon search:

- STAR Col., Phys. Rev. Lett. 114 (2015) 022301
- ALICE Col., Phys. Lett. B 752 (2016) 267


## $N \Omega$ Dibaryon

- Nucleon- $\Omega$ (N $\Omega$ ): A strangeness = -3 dibaryon is stable against strong decay
"...there is no color-magnetic effect and the energies are dominated by modification to the single-quark wave function"
- Phys. Rev. Lett. 59 (1987) 627, Phys. Rev. C69 (2004) 065207, Phys. Rev. C70 (2004) 035204.
- Scattering length, effective range and binding energy (BE) of $\mathbf{N} \Omega$-dibaryon:

|  | Scattering length <br> $\left(\mathrm{a}_{0}\right) \mathrm{fm}$ | Effective range <br> $\left(\mathrm{r}_{\text {eff }}\right) \mathrm{fm}$ | $\mathrm{BE}(\mathrm{sc})$ <br> MeV | $\mathrm{BE}(\mathrm{cc})$ <br> MeV |
| :--- | :---: | :---: | :---: | :---: |
| SU(2) | 1.87 | 0.87 | 23.2 | 19.6 |
| SU(3) | -4.23 | 2.1 | ub | ub |
| QDCSM | 2.58 | 0.9 | 8.1 | 7.3 |
| HALQCD | $-1.28+0.13^{0.14}$ | $0.499+0.026^{0.029}$ | -0.048 | $18.9+5.0^{12.1}$ |
|  |  |  |  |  |

Phys. Rev. C 83 (2011) 015202, Nucl. Phys. A 928 (2014) 89

## STAR

## Venues for Dibaryon Search

## I Systematic study of double strangeness systems



## STAR <br> N $\Omega$-dibaryon from Heavy-Ion Collisions

I $\mathrm{N} \Omega$-dibaryon is an isospin $1 / 2$ doublet and has both $\mathrm{p} \Omega$ and $\mathrm{n} \Omega$ channels possible

[ In experiments, we can look at $\mathrm{p} \Omega$ channel with two particle correlation analysis or invariant mass analysis (the $\mathrm{J}=2, \mathrm{~S}=-3$ state weakly decay is challenging)

## - Invariant mass

- Significant combinatorial background in central Au+Au collisions makes exotic particle searches difficult in heavy-ion collisions


## Two Particle Correlation in HIC



I/Baryon interaction via two particle correlation

$$
C_{A B}(Q)=\frac{N_{A B}^{\mathrm{pair}}\left(k_{A}, k_{B}\right)}{N_{A}\left(k_{A}\right) N_{B}\left(k_{B}\right)}
$$



## Lambda-Lambda Correlation Function

STAR Col. Phys. Rev. Lett. 114, 022301 (2015)


VIAll model fits to data suggest that a rather weak interaction is present between $\Lambda \Lambda$ pairs

$$
\begin{aligned}
& \text { n-n } \rightarrow \text { Phys. Lett. B, } 80(1979) 187 \\
& \text { p-n } \rightarrow \text { Phys. Rev. C 66, 047001 (2002) } \\
& \text { p-p } \rightarrow \text { Mod. Phys. } 39(1967) 584 \\
& \text { p- } \Lambda \rightarrow \text { Phys. Rev. Lett. 83, } 3138 \text { (1999) } \\
& \Lambda \Lambda \rightarrow \text { Phys. Rev. C 66, 024007(2002) } \\
& \Lambda \Lambda \rightarrow \text { Nucl. Phys. A } 707 \text { (2002) } 491
\end{aligned}
$$

## The STAR Detector at RHIC



## $\Omega$ Reconstruction (1)


$A u+A u V_{s}=200 \mathrm{GeV}$ (1.41 B events)
$\Omega \rightarrow \Lambda K$ (Mass $=1.672 \mathrm{GeV} / \mathrm{c}^{2}$ )
Branching ratio $=67.8 \%$
Mean Life time: $\tau=0.82 \times 10^{-10} \mathrm{~s}$
$\mathbf{c} \tau=2.46 \mathrm{~cm}$




Reconstructed $\Omega$


## $\Omega$ Reconstruction (2)

## Reconstructed invariant mass of $\Omega+\bar{\Omega}$






## STAR <br> Proton Identification with TPC+TOF

## Excellent PID with TPC+TOF

$\checkmark$ Number of fit points $>15$
$\checkmark$ Ratio of fit points to possible points $>0.52$
$\checkmark \mathrm{p}_{\mathrm{T}}$ cut for proton tracks $>0.15 \mathrm{GeV} / \mathrm{C}$

- DCA $<0.5 \mathrm{~cm}$
- $0.75<\mathrm{m}^{2}<1.1\left(\mathrm{GeV} / \mathrm{c}^{2}\right)^{2}$

Particle identification



With proton and anti-proton S/(S+B) ~ 99\%

## STAR Few Definitions and Corrections

Step-I Raw correlations

$$
\mathbf{C}\left(\mathbf{k}^{*}\right)=\frac{P\left(p_{a}, p_{b}\right)}{P\left(p_{a}\right) P\left(p_{b}\right)}=\frac{\text { real pairs }}{\text { mixed pairs }}
$$

$p$ - momentum of particles $a$ and $b$
Q - relative momentum
Step-II Purity correction

$$
C F_{\text {corrected }}\left(\mathrm{k}^{\star}\right)=\frac{C F_{\text {measured }}\left(\mathrm{k}^{*}\right)-1}{P P\left(\mathrm{k}^{*}\right)}+1
$$

$P P\left(k^{*}\right)=P(\Omega) \times P(p)$ is pair purity.
$P(\Omega)=S /(S+B) * F r(\Omega)$ and $P(p)=S /(S+B) * \operatorname{Fr}(p)$ where $\operatorname{Fr}(\mathrm{x})$ is Fraction of primary particles
$\operatorname{Fr}(\Omega)=1$ and $\operatorname{Fr}(\mathrm{p})=0.52$
Step-III Momentum smearing

$$
C F\left(\mathrm{k}^{*}\right)=\mathrm{CF}\left(\mathrm{k}^{*}\right) \frac{C F_{\text {nosmearing }}}{C F_{\text {smearing }}}
$$

Smearing correction factor is 0.99



## $P \Omega$ Correlation Function



PP $\rightarrow$ Pair Purity Correction
PP+SC $\rightarrow$ Pair Purity + Mom. Smearing Correction $R \rightarrow$ Emission source size
Boxes $\rightarrow$ systematic uncertainty
Comparison of measured $\mathrm{P} \Omega$ correlation function from 0-40 and 40$80 \%$ centrality with the predictions for $\mathrm{P} \Omega$ interaction potentials $\mathrm{V}_{\mathrm{l}}, \mathrm{V}_{\mathrm{II}}$ and

| Spin-2 $\mathrm{p} \Omega$ potentials | $\mathrm{V}_{\mathrm{I}}$ | $\mathrm{V}_{\mathrm{II}}$ | $\mathrm{V}_{\mathrm{III}}$ |
| :--- | :---: | :---: | :---: |
| Binding energy $\mathrm{E}_{\mathrm{B}}(\mathrm{MeV})$ | - | 6.3 | 26.9 |
| Scattering length $\mathrm{a}_{0}(\mathrm{fm})$ | -1.12 | 5.79 | 1.29 |
| Effective range $\mathrm{r}_{\text {eff }}(\mathrm{fm})$ | 1.16 | 0.96 | 0.65 | $\mathrm{V}_{\mathrm{III}}$.

[^0]
## STAR Proposal on Source Size Dependence Analysis

$\square$ The ratio of the correlation function between the small and large collision system is insensitive to the Coulomb interaction and also to the source model of the emission, thus it provides a useful measure to extract the strong interaction part of the $\mathrm{p} \Omega$ attraction from experiments at RHIC/LHC


## sTAR Source Size Analysis on $\mathrm{P} \Omega$ Correlation Function

The ratio of correlation function between small and large collision systems for the background is unity within uncertainties.

The ratio of correlation function between small and large collision systems at low $\mathrm{k}^{*}$ is lower than background.


SS $\rightarrow$ Static source ES $\rightarrow$ Expanding source
Background $\rightarrow \boldsymbol{\Omega}$ sideband is used Boxes $\rightarrow$ systematic uncertainty

| Spin-2 $\mathrm{p} \Omega$ potentials | $\mathrm{V}_{\mathrm{I}}$ | $\mathrm{V}_{\mathrm{II}}$ | $\mathrm{V}_{\text {III }}$ |  |
| :--- | :---: | :---: | :---: | :---: |
| Binding energy $\mathrm{E}_{\mathrm{B}}(\mathrm{MeV})$ | - | 6.3 | 26.9 |  |
| Scattering length $\mathrm{a}_{0}(\mathrm{fm})$ | -1.12 | 5.79 | 1.29 |  |
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| Phys. Rev. C 94, 031901 (2016) |  |  |  |  |

## Summary

V/Present the 1st measurement of correlation function for $\mathrm{P} \Omega$ from Au+Au collisions @ 200 GeV

IVThe ratio of correlation function for the small (peripheral collisions) to the large (central collisions) system is smaller than unity at low $\mathrm{k}^{*}$
$\square$ The measured ratio of correlation function from peripheral to central collisions is compared with predictions based on the $\mathrm{P} \Omega$ interaction potentials derived from lattice QCD simulations

## STAR Major Upgrades before 2020



## ViTPC Upgrade:

- Rebuilds the inner sectors of the TPC
- Continuous Coverage
- Improves dE/dx
- Extends $\eta$ coverage from 1.0 to 1.5
- Lowers $\mathrm{p}_{\text {т }}$ cut-in from $125 \mathrm{MeV} / \mathrm{c}$ to $60 \mathrm{MeV} / \mathrm{c}$


## VIEPD Upgrade:

- Allows a better and independent reaction plane measurement critical to BES physics
- Improves trigger
- Reduces background


## Status of the Inner TPC Upgrade

## - ${ }^{-1}$ SAMPA FEE (MWP2)

-2 FEEs and RDO installed on one inner most row of TPC

- Running through USB port with beam
- Design and producing pre-production RDO and FEE to instrument one Full sector for tests in fall

I Sectors (strongback + padplane + MWPC)

- Precision assembly at LBL of padplane to strongbacks and sidemounts ongoing
- Sector production started at SDU (3 completed, testing ongoing) with first fully tested sectors expected to be installed in STAR in October
[ Insertion tool
- Completed at UIC and currently being commissioned at BNL

Thank You for Your Attention!

## STARProposal on source size dependence analysis

The ratio of correlation function between small and large collision systems to extract strong p-Omega interaction w/o much contamination from Coulomb interaction.
Morita etc. arXiv:1605.06765


TABLE I: The binding energy $\left(E_{\mathrm{B}}\right)$, the scattering length $\left(a_{0}\right)$ and the effective range ( $r_{\text {eff }}$ ) with and without the Coulomb attraction in the $p \Omega$ system. Physical masses of the proton and $\Omega$ are used.

| Spin-2 $N \Omega$ Potentials |  |  |  | $V_{\mathrm{I}}$ |
| :---: | :---: | :---: | :---: | :---: |
|  | $V_{\text {II }}$ | $V_{\text {III }}$ |  |  |
|  | $E_{\mathrm{B}}[\mathrm{MeV}]$ | - | 0.05 | 24.8 |
|  | $a_{0}[\mathrm{fm}]$ | -1.0 | 23.1 | 1.60 |
|  | $r_{\text {eff }}[\mathrm{fm}]$ | 1.15 | 0.95 | 0.65 |
|  | $E_{\mathrm{B}}[\mathrm{MeV}]$ | - | 6.3 | 26.9 |
| with Coulomb | $a_{0}[\mathrm{fm}]$ | -1.12 | 5.79 | 1.29 |
|  | $r_{\text {eff }}[\mathrm{fm}]$ | 1.16 | 0.96 | 0.65 |


[^0]:    Phys. Rev. C 94, 031901 (2016)

