# The low-energy physics frontier **@Mami**: **Results and** Perspectives

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Image: SFB ≥

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#### MAMI





## Medium Energy Region





## What is it all about?



## **Scales and Phases of Nuclear Matter**

Connecting

with the

Courtesy R.F.Carsten (WNSL)



Dense Stars Astrophysics Nuclear Hypernuclei r Matter

## **First Constraint**



## QCD in the non-perturbative regime



F. Wilczek "QCD made *simple*" (<u>http://www.frankwilczek.com/</u>)

"If the Lord Almighty had consulted me before embarking upon creation, I would have recommended something *simpler*." King Alphonse X. of Castille and Léon (1221-1284), on having the Ptolemaic system of epicycles explained to him

#### **First Constraint**



Modern nuclear physics is about...

## →Linking QCD to many body systems



Universal Nuclear Energy Density Functional

## The Beauty of the Electromagnetic Probe

Clean probe of hadron structure

- > Electron-vertex wellknown from QED
- > One-photon exchange dominates
- > Higher-order exchange diagrams are suppressed



> Vary the wavelength of the probe to view deeper inside the hadron

One Millennium Quest Generation of Mass



#### Not the elementary mass of the fermions $\rightarrow$ Higgs Sector

But the actual mass of the "Macroscopic" Hadron and its Composites



## Nuclear Physics @ A=1 = Nucleon Properties

#### mass

 $m_{
m p} = 938.272046(21)\,{
m MeV/c^2}$ Discovered by E. Rutherford (1919)

 $m_n = 939.565379(21) \,\mathrm{MeV/c^2}$ 

Discovered by J. Chadwick (1939)

#### size

moments of electric charge and magnetization distribution derived from **form factor** measurements



## **Form Factors are Ethernal**

#### http://hyperphysics.phy-astr.gsu.edu/



#### Form Factors from Elastic ep scattering

form factor: 
$$G(q^2) = \frac{1}{e} \int_0^\infty \rho(r) \frac{\sin qr}{qr} 4\pi r^2 dr$$



#### **The latest MAMI measurement**

## The experiment designed for ... high precision by redundancy

PRE 105, 242001 (2010) PHYSICAL REVIEW LETTERS 10 DECEMBER 2010 Figh-Precision Determination of the Electric and Magnetic Form Factors of the Proton J.C. Bernauer,<sup>1,\*</sup> P. Achenbach,<sup>1</sup> C. Ayerbe Gayoso,<sup>1</sup> R. Böhm,<sup>1</sup> D. Bosnar,<sup>2</sup> L. Debenjak,<sup>3</sup> M.O. Distler,<sup>1,\*</sup> L. Doria,<sup>1</sup> A. Esser,<sup>1</sup> H. Fonvieille,<sup>4</sup> J. M. Friedrich,<sup>5</sup> J. Friedrich,<sup>1</sup> M. Gómez Rodríguez de la Paz,<sup>\*</sup> M. Makek,<sup>2</sup> H. Merkel,<sup>1</sup> D.G. Middleton,<sup>1</sup> U. Müller,<sup>1</sup> L. Nungesser,<sup>1</sup> J. Pochodzalla,<sup>\*</sup> M. Potokar,<sup>5</sup> S. Sánchez Majos,<sup>\*</sup> B.S. Schlimme,<sup>1</sup> S. Širca,<sup>6,5</sup> Th. Walcher,<sup>1</sup> and M. Weinriefer<sup>1</sup>

- Statistical precision  $\sigma < 0.1\%$
- $\delta\theta < 0.5$  mrad vertical and horizontal
- Control of luminosity and systematic errors

#### All quantities measured by more than one method

#### **Rosenbluth with a twist**

#### "Super-Rosenbluth Separation": fit of form factor models **DIRECTLY to cross sections**



- All Q<sup>2</sup> and ε data are used in one fit
- No projection to constant Q<sup>2</sup>
   no limit of kinematics
- One "estimator"
  - → stat. theory "robust estimator"





#### **The Low Energy Frontier**





→ Radius of proton is dominant uncertainty in many QED processes



0.8

0.85

0.9

0.95

proton charge radius (fm)

**muonic and electronic measurements** (atomic)

6

Delayed / prompt events (10-4)

3

0

49.75





**Discrepancy is between muonic and electronic measurements** (scattering)



#### Novel beyond SM physics? Novel hadron physics?

already excluded: missing atomic physics, structures in FF, anomalous 3rd Zemach radius





#### **Discrepancy is between muonic and electronic measurements** (both types)

New data are needed and they are coming ...

#### **Improvements on Proton Radius**

Possible experiments include:

Redoing atomic hydrogen

Light muonic atoms for radius comparison in heavier systems





#### **Second constraint**

© Jens Rydén



#### **Second constraint**

## It's always 50 just nuclear matter

#### **Second constraint**



Modern nuclear physics is about...

## →Unravelling the phases of nuclear matter



LRP Nuclear Science Advisory Committee(2008)

## The Equation of State of Nuclear Matter

A heavy nucleus (like <sup>208</sup>Pb) is 18 orders of magnitude smaller and 55 orders of magnitude lighter than a neutron star

Yet bounded by the same **EOS** 





#### **The Low Energy Frontier**

#### Strong Interactions Hadron Structure Hadron Spectroscopy

N<sub>skin</sub> Heavy Nucleus





...the Astrophysics frontier

**High-Energy Physics** 

**Precision Physics** 

**Atomic Physics** 

**Astrophysics** 

#### **Neutron Skin Measurements**

#### Where do the neutrons go?



#### Pressure forces neutrons out against surface tension





#### **Neutron Skin Measurements**

#### Where do the neutrons go?



#### Pressure forces neutrons out against surface tension





#### PV: the view in the mirror



Electron scattering  $\gamma$  exchange provides R<sub>p</sub> through nucleus FFs

PV e-scattering

Electron also exchange Z, which is parity violating

Primarily couples to neutron

		•••
electric charge	1	0
weak charge	≈0.07	1

#### PV: the view in the mirror



Electron scattering  $\gamma$  exchange provides  $R_p$  through nucleus FFs

PV e-scattering Electron also exchange Z, which is parity violating Primarily couples to neutron

Detectable in PV asymmetry of electrons with different helicity



#### Hard, harder, PV Experiments



## Neutron skin Radii: Where are we?



## Diverse experiments but consistent results

Solution Too many model dependent observables

#### Precise determination of N<sub>Skin</sub> in <sup>208</sup>Pb set a basic constraint on the nuclear symmetry energy

X. Roca-Maza, at al. Phys. Rev. Lett. 106, 252501 (2011)



#### At A2 we shine light on the nucleus!

coherent  $\pi^0$  photoproduction @ A2

$$\begin{array}{c} & \overbrace{} & \overbrace{} & \overbrace{} & \overbrace{} & \overbrace{} & \overbrace{} & \gamma \\ & & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\$$

# $\pi^0$ produced with equal probability on protons 0 and neutrons 0

**Most simple** - **P**lane Wave Impulse Approximation:

$$\frac{d\sigma}{d\Omega}(\mathbf{PWIA}) = \frac{s}{m_N^2} \times \frac{1}{2} \frac{q_\pi^*}{k^*} \left| F_2\left(E_\gamma^*, \theta_\pi^*\right) \right|^2 \sin^2\left(\theta_\pi^*\right) \times A^2 F^2\left(q\right)$$

#### Matter Form Factor $\rightarrow$ r.m.s. matter radius

## $^{208}\text{Pb}$ neutron skin from Coherent $\pi^{\text{o}}$



D. Watts, et al, EPJ Web of Conferences 37, 01027 (2012)





New experimental campaign@A2 (October 2012) 116,120,124Sn, <sup>58</sup>Ni, <sup>208</sup>Pb

#### **The Low Energy Frontier**



#### **Conventional strategies for DM searches**

#### A **bottom up** approach: Looking for interacting particles

#### SM X SM X

Direct Production:

Tevatron, LHC



Direct Search:

CDMS, DAMA/LIBRA, XENON, CRESST, LUX, COUPP, KIMS, ...

Indirect Search:

PAMELA, Fermi, HESS, ATIC, WMAP, ...



#### **Conventional strategies for DM searches**

#### A **bottom up** approach: Looking for interacting particles

#### Assumptions:

X

χ

X

SM

SM

SM

SM

SM

χ

SM

χ

χ

There is dark matter (SUSY or something else) Dark matter interacts with Standard Model matter (besides gravity) Dark matter interacts via a "dark force"

#### **Question:**

What is the character of this "dark force"? Scalar, pseudo-scalar, vector bosons? Massive or mass-less? Mass range? Size of the coupling constant?

## A top down motivation

- Extra U(1) gauge bosons ubiquitous in extension of SM
- U(1) gauge bosons may be hidden (no interaction with SM)
- No reason for U(1) boson to be heavy
- Dark matter couples to U(1) bosons y and y'



- Mixing parameter  $\epsilon$  of  $\gamma/\gamma'$  mixing
- Boson mass  $m_{Y'} > 0 \Rightarrow$  decay suppressed, macroscopic lifetime

 $\Rightarrow$  Look for  $\chi$  at high energies OR for  $\gamma'$  at low energies!

B. Holdom Phys. Lett. B 166 (1986) 196

#### **Probing Dark Forces @ GeV Scale**

## Dark Photon

#### Light weakly coupled U(1) gauge boson

N. Arkani-Hamed, et al., Phys. Rev. D 79 (2009) 015014

...it explains ... terrestrial anomalies (DAMA, CDMS, XENON) satellite anomalies (PAMELA, FERMI)

(g-2)<sub>µ</sub> anomaly M. Pospelov, Phys. Rev. D80 (2009) 095002

#### **Proton Radius Puzzle**

D. Tucker-Smith and I. Yavin Phys. Rev. D83 (2011) 101702

#### PHYSICAL REVIEW D 80, 075018 (2009)

#### New fixed-target experiments to search for dark gauge forces

James D. Bjorken,<sup>1</sup> Rouven Essig,<sup>1</sup> Philip Schuster,<sup>1</sup> and Natalia Toro<sup>2</sup> <sup>1</sup>Theory Group, SLAC National Accelerator Laboratory, Menlo Park, California 94025, USA <sup>2</sup>Theory Group, Stanford University, Stanford, California 94305, USA (Received 20 July 2009; published 28 October 2009)



World wide effort (CERN, DESY, JLAB, MAMI, all e<sup>+</sup>e<sup>-</sup> colliders, ...)

#### Prediction are testable: Large cross section in leptons

## Search for the Dark Photon @ MAMI

H. Merkel et al., Phys. Rev. Lett. 106 (2011) 251802

#### **Bump Hunt**: Quasi-photoproduction off <sup>181</sup>Ta target



Beam current: 100µA Luminosity: L = 1.7.10<sup>35</sup> (s.cm<sup>2</sup>)<sup>-1</sup> Minimal angles for spectrometers Geometry as symmetric as possible (background reduction)



## Search for the Dark Photon @ MAMI

H. Merkel et al., Phys. Rev. Lett. 106 (2011) 251802

#### **Bump Hunt**: Quasi-photoproduction off <sup>181</sup>Ta target

New E/my scan



**Future**: Low mass region <50 MeV/ $c^2$  and small dark photon coupling  $\epsilon^2$ 

## ... MESA beyond MAMI ...

#### Mainz Energy recovering Superconducting Accelerator



- Superconducting LINAC
- Three recirculation arcs for external beam high external current for "next generation" parity violation experiments
- Energy recovery mode (half wave-length recirculation) for *internal target* experiments

	Current	Energy	Luminosity
External Beam Mode:	150 <i>μ</i> Α	200 MeV	$10^{39}{ m cm}^{-2}{ m s}^{-1}$
Energy Recovery Mode:	10 mA	150 MeV	$10^{36}{ m cm}^{-2}{ m s}^{-1}$



#### The Low Energy Frontier @ Mainz: Results and ...





#### ... perspective are excellent!

Neutron form factor: new neutron detector Nucleon polarizabilities with real photons High resolution pion spectroscopy of light hypernuclei

