Retardation effect for heavy parton collisional energy loss in QGP.

- Motivation
- Collisional energy loss in stationnary regime
- Retardation effect; main results
- Radiation
- Conclusion & Perspectives

Pol-Bernard Gossiaux work done with Stéphane Peigné & Thierry Gousset SUBATECH (Nantes) hep-ph/0509185

...collisional energy loss...

Recent papers advocating the possible role of this mechanism for partons of moderate energy: Dutt Mazumder et al (Phys.Rev. D71 (2005) 094016) / Mustafa et al (Phys.Rev. C72 (2005) 014905)/ Gay Ducati et al (hep-ph/0506241) /

But not the main stream



due to colinear divergences

Usual QGP: no quasi bound state, at rest, static, « confined » over some extansion L_{plasma} ... (Boring) heavy parton collisional energy loss in QGP.

Nearly everything is known; Braaten & Thoma (91)





Figure taken from Mustafa et al (Phys.Rev. C72 (2005) 014905)



Physically reasonable, but HOW BIG?

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CEL in stationnary regime.

Assume pQCD regime at high temperature: g<<1 and ordering

$$\frac{1}{M} << r = \frac{1}{T} << \frac{1}{q^*} << m \approx \frac{1}{gT} << \lambda \approx \frac{1}{g^2 T}$$



Heavy quark probes the medium via virtual gluon of momentum *k*

Zone I: *k*>*q** : hard; close collisions; individual; incoherent.

-Zone II: *k*<q*: soft; far collisions; collective; coherent; macroscopic.

CEL in stationnary regime: close collisions.

Bjorken (82) evaluated this contribution using kinetic theory.



From rate to energy loss :

$$\frac{dE}{dx} = \int \frac{d\Gamma(E',E)}{dE'} \times \frac{E'-E}{v} dE' = \frac{2}{3} \alpha m_D^2 \ln\left(\frac{\sqrt{ET}}{q^*}\right) \quad \text{for } v \approx 1.$$
IR divergent

CEL in stationnary regime: far collisions.

Will embed most of the retardation effect.

First evaluated in stationnary regime by Braaten & Thoma (82) following Weldon (83): $\Gamma = -(1-n(E))/2E \times tr[(P.\gamma+M) Im(\Sigma)]$



Quite general relation : includes the soft collisions (k<q*) as well, provided one takes the full gluon propagator Π :

CEL in stationnary regime: far collisions (II).



For k>q*: through kinetic expressions

For k<q*: evaluate the self energy Σ with HTL ressumation for Π (consistent with collective polarization)... Imaginary time formalism.

Legitimate as both contributions are separately gauge invariant.

CEL in stationnary regime: far collisions (III).

Results at the logarithmic accuracy:



Far collisions: alternative formulation

Maxwell's Equations with appropriate dielectric functions (collective response of the medium in linear response approx.) lead to the same result for Energy loss.

Easy strategy: evaluate induced chromoelectric field (medium polarization Eind and the Lorentz force of this field on the heavy parton / external current)

© Easy to generalize for non stationnary phenomena (heavy parton produced inside the QGP)... better suited to discuss physics.

 \otimes Misses some part of the hard dynamics... which plays little role for retardation \otimes .

N.B.: L.R. \Rightarrow trivial color structure of dielectric functions.

Far collisions: alternative formulation (II)

Maxwell equations in Fourier space $(\omega, k) \Rightarrow$

$$\mathcal{E}_L \vec{E}_L + (\mathcal{E}_T - k^2 / \omega^2) \vec{E}_T = \frac{4\pi}{i\omega} (\vec{j}_L + \vec{j}_T)$$

With εL and εT : dielectric functions of the plasma

$$\varepsilon_L(k,\omega) = 1 + \prod_L(\omega/k)/k^2$$
; $\varepsilon_T(k,\omega) = 1 - \prod_T(\omega/k)/\omega^2$

(Polarization Π taken in H.T.L approximation)

 $\vec{j}_{L}(\vec{j}_{T})$: longitudinal (transverse) projection of the current on \vec{k}

$$\vec{E}_{ind}(\vec{k},\omega) = \frac{4\pi}{i} \left[\frac{k^2}{k^2 + \Pi_L} \times \frac{\vec{j}_L}{\omega} + \frac{\omega}{\omega^2 - k^2 + \Pi_T} \times \vec{j}_T \right]_{ind} = 4i\pi \left[k^2 \Delta_L \times \frac{\vec{j}_L}{\omega} + \omega \Delta_T \times \vec{j}_T \right]_{ind}$$
Substract the vacuum value gluon propagator HTL

Far collisions: alternative formulation (III)

Fourier back and integrate the Lorentz force up to t=L/v:

$$-\Delta E(L) = i \sum_{a} \int \frac{d^{3}k}{4\pi^{3}} \int_{-\infty}^{+\infty} d\omega \left[\frac{k^{2}}{k^{2} + \Pi_{L}} \frac{\vec{j}_{L,a}(K)}{\omega} + \frac{\omega}{\omega^{2} - k^{2} + \Pi_{T}} \vec{j}_{T,a}(K) \right]_{ind} \cdot \int_{0}^{L/v} dt \ e^{-iK\cdot V t} \ q_{a} \ \vec{v}(t)$$

Special case of stationnary current: $j_{a}^{\mu}(t,\vec{x}) = q_{a}V^{\mu}\delta^{3}(\vec{x}-\vec{v}t)$ with $V^{\mu} = (1,\vec{v})$
$$\Rightarrow \quad j_{a}^{\mu}(K) = 2\pi \ q_{a}V^{\mu} \ \delta(K\cdot V) = 2\pi \ q_{a}V^{\mu} \ \delta(\omega-\vec{k}\cdot\vec{v}) = L/V$$

$$\frac{-\Delta E(L)}{C_{R}\alpha_{S}L} = \underbrace{i}_{V} \int \frac{d^{3}k}{2\pi^{2}} \int_{-\infty}^{+\infty} d\omega \left[\frac{k^{2}}{k^{2} + \Pi_{L}} \frac{v_{L}^{2}(K)}{\omega} + \frac{\omega}{\omega^{2} - k^{2} + \Pi_{T}} v_{T}^{2}(K) \right]_{ind} \delta(\omega-\vec{k}\cdot\vec{v})$$

Only imaginary part contributes

N.B.: « induced » not really needed

Far collisions: alternative formulation (IV)



In stationnary regime: Only the $Im(\varepsilon)$ contributes (ok); collective modes do not get excited; no Cherenkov radiation.

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Retardation effect: model for partonic current



For partons produced at t=0

$$j_0^{\mu a}(K) = iq^a \left(\frac{V_1^{\mu}}{K \cdot V_1 + i\eta} - \frac{V_2^{\mu}}{K \cdot V_2 + i\eta} \right)$$

with $V_1 = (1, \vec{v}_1)$ and $V_2 = (1, \vec{v}_2)$

Satisfies current conservation $K \cdot V = 0$

Simplifying assumption: $v_2=0 \Rightarrow V_2=(1,0)$ and $V_1=(1, \vec{v})$

Only **one** parton contributes to \vec{j} , but the system is in fact color neutral.

Retardation effect: generalization of ΔE

$$-\Delta E(L) = i \sum_{a} \int \frac{d^{3}k}{4\pi^{3}} \int_{-\infty}^{+\infty} d\omega \left[\frac{k^{2}}{k^{2} + \Pi_{L}} \frac{j_{Ld}(K)}{\omega} + \frac{\omega}{\omega^{2} - k^{2} + \Pi_{T}} \frac{j}{p_{d}(K)} \right]_{ind} \cdot q_{d} \vec{v} \int_{0}^{V} \frac{d}{\psi} e^{i(KVI)} KVI = 0$$

$$\frac{-\Delta E(L)}{C_{R}\alpha_{s}} = -iv^{2} \int \frac{d^{3}\vec{k}}{4\pi^{3}} \int_{-\infty}^{\infty} \frac{d\omega}{\omega} \left[k^{2} \cos^{2} \theta \Delta_{L}(\omega, k) + \omega^{2} \sin^{2} \theta \Delta_{T}(\omega, k) \right]_{ind}$$

$$complex \longrightarrow \left\{ \begin{array}{c} 1 - e^{-i(\omega - kv \cos \theta) L/v} \\ (\omega - kv \cos \theta) & \omega - kv \cos \theta + i\eta \end{array} \right\}$$

Cut rent
Poles Coll modes Coll modes Coll modes Coll modes Coll modes Coll modes Coll mode Col

Retardation effect:convergence and scales

Contour integration in lower half-(complex) plane :

$$\frac{-\Delta E(L)}{C_R \alpha_s} = \int \frac{d^3 \vec{k}}{2\pi^2} \left\{ \frac{1 - e^{ikL\cos\theta}}{k^2 + m_D^2} + v^2 \int_{-\infty}^{\infty} \frac{d\omega}{2\pi\omega} \left[k^2 \cos^2\theta \,\rho_L + \omega^2 \sin^2\theta \,\rho_T \right] \right. \\ \left. \times \operatorname{P}\left(\frac{1}{\omega - kv\cos\theta} \right) \frac{1 - e^{-i(\omega - kv\cos\theta)L/v}}{\omega - kv\cos\theta} \right\}_{\mathrm{ind}}.$$

With the spectral functions :

$$\rho_s(\omega, k) \equiv 2 \operatorname{Im} \Delta_s(\omega + i\eta, k)$$

= $2\pi \operatorname{sgn}(\omega) z_s(k) \delta(\omega^2 - \omega_s^2(k)) + \beta_s(\omega, k) \theta(k^2 - \omega^2)$

Residue Magnitude of the cut

Retardation effect: convergence and scales

Contour integration in lower half-(complex) plane :

$$\frac{-\Delta E(L)}{C_R \alpha_s} = \int \frac{d^3 \vec{k}}{2\pi^2} \left\{ \underbrace{\frac{1 - e^{ikL\cos\theta}}{k^2 + m_D^2}}_{\sum k^2 + m_D^2} + v^2 \int_{-\infty}^{\infty} \frac{d\omega}{2\pi\omega} \left[k^2 \cos^2\theta \rho_L + \omega^2 \sin^2\theta \rho_T \right] \right. \\ \left. \left. \left. \times \mathbf{P} \left(\frac{1}{\omega - kv\cos\theta} \right) \frac{1 - e^{-i(\omega - kv\cos\theta)L/v}}{\omega - kv\cos\theta} \right\}_{\text{ind}} \right\}_{\text{ind}} .$$

 \rightarrow finite constant when L $\rightarrow \infty$ (due to the induced prescription)

Retardation effect: convergence and scales

Contour integration in lower half-(complex) plane :

$$\frac{-\Delta E(L)}{C_R \alpha_s} = \int \frac{d^3 \vec{k}}{2\pi^2} \left\{ \frac{1 - e^{ikL\cos\theta}}{k^2 + m_D^2} + v^2 \int_{-\infty}^{\infty} \frac{d\omega}{2\pi\omega} \left[k^2 \cos^2\theta \rho_L + \omega^2 \sin^2\theta \rho_T \right] \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{w^2 + m_D^2} + v^2 \int_{-\infty}^{\infty} \frac{d\omega}{2\pi\omega} \left[k^2 \cos^2\theta \rho_L + \omega^2 \sin^2\theta \rho_T \right] \right\}_{\text{ind}} \cdot \frac{\Delta E(L)}{v} = \int_{k < k_{cut}} \frac{d^3 \vec{k}}{2\pi^2} \left\{ \frac{1 - e^{ikL\cos\theta}}{k^2 + m_D^2} + v^2 \int_{-\infty}^{\infty} \frac{d\omega}{2\pi\omega} \left[k^2 \cos^2\theta \rho_L + \omega^2 \sin^2\theta \rho_T \right] \right\}_{\text{ind}} \cdot \left[\left(2 \frac{\sin^2 \left((\omega - kv \cos\theta) L/(2v) \right)}{(\omega - kv \cos\theta)^2} - \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right) + \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right] \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{v^2 + m_D^2} + \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} - \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{v^2 + m_D^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} - \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{v^2 + m_D^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} - \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{v^2 + m_D^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} - \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{v^2 + m_D^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} - \frac{\pi L}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v} \delta(\omega - kv \cos\theta) \right\}_{\text{ind}} \cdot \frac{1}{v} \left\{ \frac{1 - e^{ikL\cos\theta}}{(\omega - kv \cos\theta)^2} + \frac{1}{v$$

Brings a 1/k additional factor instead of L \Rightarrow UV convergent !

Exactly the stationnary E loss log divergent in UV

Retardation effect: convergence and scales (II)

<u>Conclusion</u>: $-\Delta E(L) = dE_{stat}/dx * L + B(L)$

With: B(L) UV safe (as anticipated) and \rightarrow cst when L $\rightarrow \infty$.

For small L, the hard scale $\ll 1/L \gg$ would lead to k_{typ} in B(L) above $k_{cut} \Rightarrow$ Condition on L: must be typically larger then $1/m_D$.

Behaviour at small L has been nevertheless studied, using *sum rules*; interesting cancelations were discovered.

$$-\Delta E(L) = \frac{C_R \alpha_s m_D^2 k_{cut}^3}{108 \, \pi v^2} L^4$$

Retardation effect: Results





Retardation effect: Results (II)

 $-\Delta E_{asymp}(L,E,\ldots) = dE_{stat}/dx * (L-delay(E,T))$



Delay roughly α 1/T, but not of the order of rD !!!

Retardation effect: Results (III)

Might have some phenomenological implications :

Jet absorption and corona effect at RHIC. Extracting collision

geometry from experimental data

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Abstract

We propose a simple model based on Monte Carlo simulation of nucleus-nucleus collisions using a Glauber approach to explain experimental data on the angular dependence of the nuclear modification factor R_{AA} at high transverse momentum in the reaction plane. The model has one free parameter $L\simeq 2$ fm to describe the the thickness of the corona area and was ajusted to fit the experimental data on AuAu collisions at centrality 50-60%. The model nicely describes the R_{AA} dependence for all centrality classes. We extract the second Fourier component amplitude, v_2 , for high pt particle azimuthal distribution and found v_2 should be at the level of 11-12% purely from the geometry of the collision with particle absorption in the core. We give a prediction for R_{AA} in Cu+Cu collisions at 200 GeV. Our physical interpretation of the parameter L is that it's actually the formation time $T = L/c \simeq 2$ fm/c.

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Retardation effect... and other related mechanisms.



 \rightarrow finite negative constant when L $\rightarrow \infty$ (dipole separation easier in QGP than in vacuum)

Retardation effect... and other related mechanisms.

- If one just considers the force acting on the heavy parton: irrelevant question (It takes some time to reach its fully value, that's all).
- Being less inclusive, one can identify another interesting mechanism...

Retardation effect... and radiation.



 $\delta(\omega - \omega_s(k))$: • Clear signature of some radiation (cf. identification of Cherenkov radiation in stationnary Collisional E loss)

• Does not scale like $L \Rightarrow$ Not Cherenkov (although emitted by the time dependent medium polarization), but 'initial' Brehmstrallung .

• ΔE incorporates the difference of energies radiated in the medium (neglecting radiative rescattering) and in the vacuum.

In-vacuum radiation



Radiative contribution W(L) to the energy loss :

$$\frac{dW(L)}{dk\,d\cos\theta}\Big|_{vac} = \frac{C_R\alpha_s}{\pi}\sin^2\theta\,\frac{\sin^2((k-kv\cos\theta)\,L/(2v))}{(\cos\theta-1/v)^2} \quad \xrightarrow{}_{L\to\infty}\frac{C_R\alpha_s}{2\pi}\frac{v^2\sin^2\theta}{(1-v\cos\theta)^2}$$

<u>N.B.</u>: For v=1, this is just the Z.O.L. result of GLV for x<<1

"Radiates like crazy", but what should matter is the difference with the "In-medium" radiation.

In-medium radiation



Radiation of collective modes, described at the level of the HTL ressumation

Not considered before, to my knowledge

$$\frac{dW(L)}{dk\,d\cos\theta} = \frac{C_R\alpha_s}{\pi} \left\{ z_L(k) \frac{k^2}{\omega_L^2(k)} \cos^2\theta \frac{\sin^2((\omega_L(k) - kv\cos\theta)L/(2v))}{(\cos\theta - \omega_L(k)/(kv))^2} + z_T(k)\sin^2\theta \frac{\sin^2((\omega_T(k) - kv\cos\theta)L/(2v))}{(\cos\theta - \omega_T(k)/(kv))^2} \right\}$$
$$\xrightarrow{L\to\infty} \frac{C_R\alpha_s}{2\pi} \left\{ z_L(k) \frac{k^2}{\omega_L^2(k)} \frac{\cos^2\theta}{(\cos\theta - \omega_L(k)/(kv))^2} + z_T(k) \frac{\sin^2\theta}{(\cos\theta - \omega_T(k)/(kv))^2} \right\}$$

Long: suppressed at large k

Vacuum vs Infinite medium







• In all cases, the θ -integrated Inmedium radiation is weaker than in the vacuum (finite mass, residues < 1)

• Interplay of trans. and long. radiation for ω of the order of mD

Infinite medium vs Finite (L=5fm) medium







- Depletion at small angle
- <u>Gluon formed</u>:

 $\sin^2((\omega_L(k) - kv\cos\theta)L/(2v))$ averages to $\frac{1}{2}$

Provides a criteria for formation time.

• Possible consequences on the understanding of far away hadrons.

Conclusions & Perspectives.

- 1. Extansion of the formalism of Thoma and Guylassy in the case of partons produced inside the medium.
- 2. Collisional energy loss suppressed by large factor. Retardation time of the order of several fm/c.
- 3. In-vacuum gluon radiation is supressed by the medium. Radiation of collective mode is treated using the correct dispersion relation.
- 4. Crucial implications on the phenomenology of jet quenching.
- 5. Consequences on the radiative energy loss induced by rescatterings ?



the splitting function of some in-medium evolution equation ?